

Gravity in the long range.

H/W

Review the concept of natural units used in cosmology & high energy physics.
→ 4 fundamental units, their conversion from SI to natural units.

Ref.: Chapter 3 - B. Ryden

→ The fundamental forces in nature

- * Gravity
- * Electromagnetism
- * Weak nuclear force
- * Strong nuclear force.

④ Which one dominates at short length scale?
— at long " " ?

→ Strong and weak nuclear forces dominate in the lengthscale $\lesssim 10^{-15}$ m.

→ The observable universe is mostly neutral \Rightarrow negligible electromagnetic interaction.

\Rightarrow Gravity is the only dominant force !!
(being the weakest at small length scale).

② Some interesting numerical calculations:
 → what if each atom on earth would be missing an e^- ?
 → ~~between two e^- / p^+ p^+ , e^- at what distance gravity would~~

→ Gravity — Newtonian or Einsteinian?

* At least, in our solar system, we as per the Newtonian gravity any massive object experiences a force due to another mass, given by

$$F = - \frac{G M_g m_g}{r^2}$$

where M_g, m_g are respective masses of obj. that play a part in gravitation.

→ But, as per Newtonian mechanics,

$$F = \underbrace{m_i}_{\substack{\text{let's call this } m_i \\ \uparrow \\ \text{inertial mass.}}} a$$

② Is m_i the same as m_g ?

→ YES!!! due to Galileo!

In principle,

Given a gravitating mass M_g , an object with inertial mass m_i would move as per acceleration

$$a = -\frac{GM_g}{r^2} \left(\frac{m_g}{m_i} \right)$$

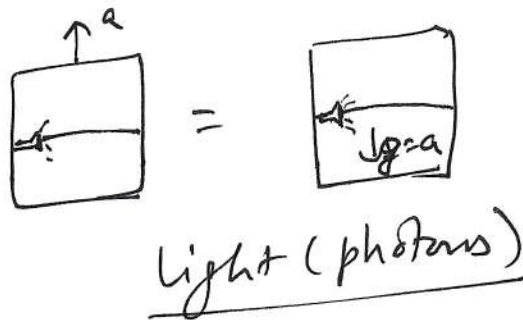
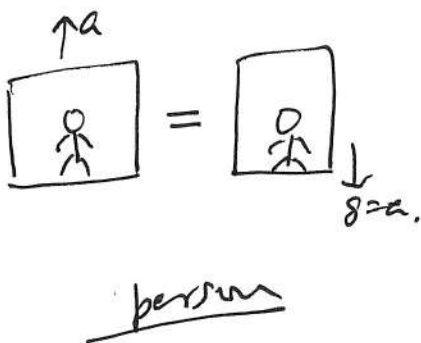
→ But we know that all objects fall to earth's surface with same acceleration!

⇒

$$m_g = m_i$$

Equivalence principle.

The basis for Einstein's theory of gravitation.



→ But light has no inertial mass!!
→ Also, acc. to ~~for~~ Fermat's principle in optics, sp. light takes the shortest route
↳ In flat space, it'd be a straight line. ↳ Euclidean.

→ But, in presence of gravity if light is curved
⇒ the space can't be Euclidean!

↓
A curved space(-time)?

Newton

* Mass tells gravity how to exert force

* Force tells mass how to move.

Einstein.

* Mass-energy tells spacetime how to curve.

* Curvature tells ~~mass~~ mass-energy how to move.
(geodesic)

⇒ Natural explanation of equivalence principle.

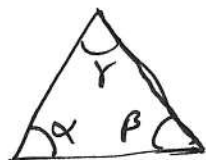
⑨ So how do we describe this curvature?

→ Before going to STR/GTR, let's consider a 2D space.

→ We know, from trigonometry

$$\alpha + \beta + \gamma = \pi$$

↳ valid for a plane.



→ The distance between two infinitesimally separated pts.

$$ds^2 = dx^2 + dy^2 \\ = dr^2 + r^2 d\theta^2$$

Why if we draw Δ^u in sphere

$$\alpha + \beta + \gamma = \pi + \frac{A}{R^2} \leftarrow \text{+ve curvature.}$$

→ on a saddle :

$$\alpha + \beta + \gamma = \pi - \frac{A}{R^2} \leftarrow \text{-ve curvature.}$$

& the distance bet. pts :

$$ds^2 = dr^2 + R^2 \sin^2\left(\frac{r}{R}\right) d\theta^2 \leftarrow \text{+ve}$$

$$ds^2 = dr^2 + R^2 \sinh^2\left(\frac{r}{R}\right) d\theta^2 \leftarrow \text{-ve.}$$

→ In 3D:

$$ds^2 = dr^2 + S_\kappa(r)^2 [d\theta^2 + \sin^2\theta d\phi^2]$$

$$\text{where, } S_\kappa(r) = \begin{cases} R \sin\left(\frac{r}{R}\right) & , \kappa = +1 \\ r & , \kappa = 0 \\ R \sinh\left(\frac{r}{R}\right) & , \kappa = -1 \end{cases}$$

A Brief review of STR [Discussion > Lecture]

→ From the last lecture:

The equivalence principle implies that the geometry of "space" cannot be Euclidean!

Ⓞ Then what can it be?

→ To answer this question, we must revert to the "local" (small) lengthscale instead of large lengthscale.

→ Since it's the geometry of frames of references we are interested in, we consider free space (free of interactions).

→ Historically, this discussion ^{STR} precedes the discussion about gravity _{GTR}.

→ We start with one of the most fundamental (earliest) principles of physics:

* The principle of relativity (by Galileo)

"No expt. can measure the absolute vel. of an observer; results of expt. are independent of the relative vel. of observer w.r.t others."

⇒ Galilean geometry.

→ Einstein's contribution:

"The universality of the speed of light."
i.e. two unaccelerated observers will measure the same speed of light.

⑧ How does this "postulate" change/modify the principle of relativity?

Ex: In Galilean relativity, relative speed of light:

$$c' = c - v \quad \leftarrow \text{Not true!}$$

→ Clearly, one would have to a new (mathematical) coordinate sys.!

→ Define an inertial frame in "special" relativity:

- (1) distance between two spatial pts. is independent of time.
- (2) All clocks in a frame are synchronized. & run at same rate.
- (3) At a given time t , geometry is Euclidean.

→ Units:

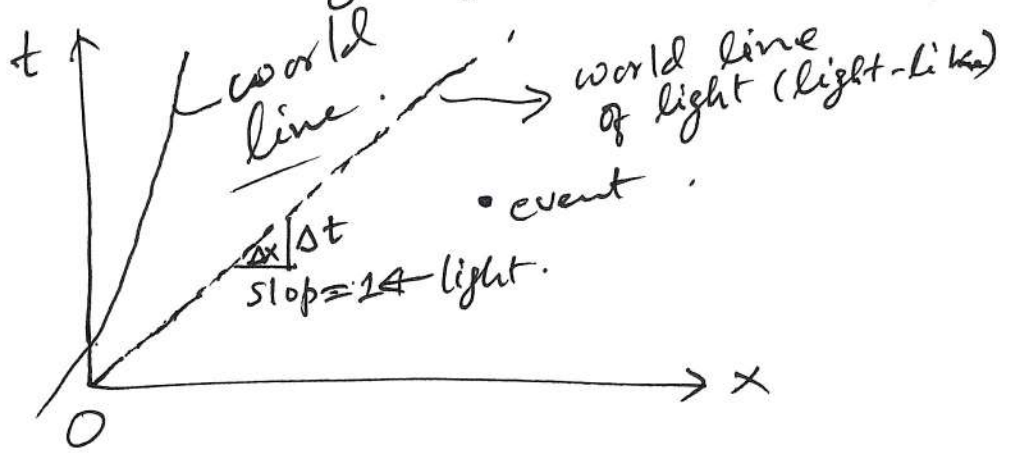
We redefine length & time SI units, so that

$$3 \times 10^8 \text{ ms}^{-1} = 1$$

$$\Rightarrow 1 \text{ s} = 3 \times 10^8 \text{ m}$$

$$\Rightarrow \boxed{c = 1} \quad \leftarrow \text{dimensionless \& unity.}$$

→ A world line = diag. of $x-t$!



→ Recall: The universality of light implies that the ~~speed of~~ $c = 1$ for all ~~obs~~ inertial observers (O, O', \bar{O}, \dots).
~~movin~~

In some observer's frame O :

$$\frac{\Delta x}{\Delta t} = 1$$

$$\Rightarrow \Delta x = \Delta t$$

$$\Rightarrow \Delta x^2 = \Delta t^2 \quad \leftarrow \text{in terms of distances}$$

$$\Rightarrow \Delta x^2 + \Delta y^2 + \Delta z^2 = \Delta t^2 \quad \leftarrow \text{Going to 3D}$$

$$\Rightarrow \boxed{\Delta x^2 + \Delta y^2 + \Delta z^2 - \Delta t^2 = 0}$$

Why in frame O' :

$$\Delta x'^2 + \Delta y'^2 + \Delta z'^2 - \Delta t'^2 = 0$$

→ Here, t & t' must be different because simultaneity of ~~it~~ is not guaranteed between two frames.

$$\Rightarrow \Delta x'^2 + \Delta y'^2 + \Delta z'^2 - \Delta t'^2 = \Delta x^2 + \Delta y^2 + \Delta z^2 - \Delta t^2$$

H/W Read Schutz Sec. 1.6 for the proof.

$$\left[\equiv \Delta s^2 \right]$$

"The invariant" of special relativity.

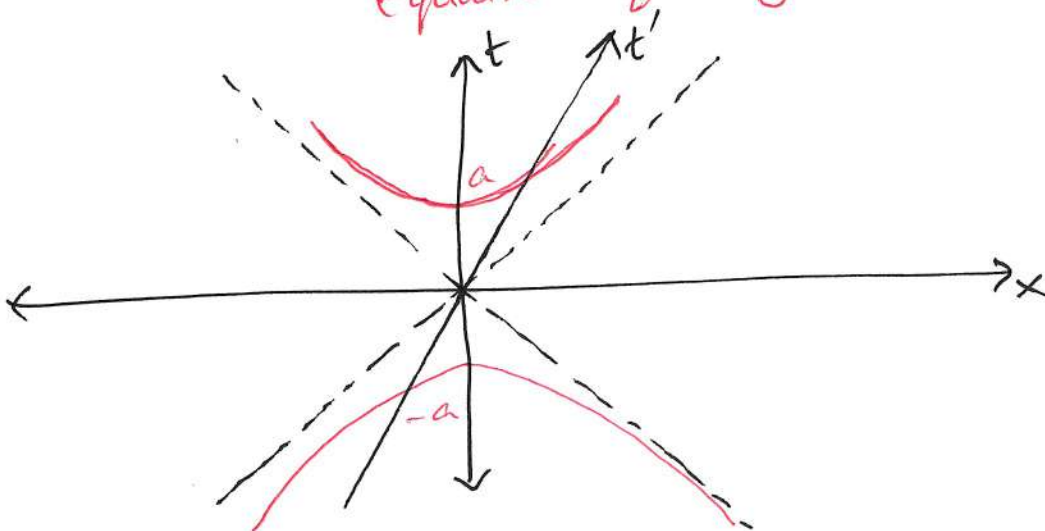
NOTE: This Δs^2 is unlike the invariant of Euclidean geometry which would have been $\Delta x_i^2 + \Delta t^2$!!

→ The space-time curve parametrized by the invariant Δs^2 :

Consider 1+1D space (Δ time)

i.e.
$$-t^2 + x^2 = a^2$$

Equation of a hyperbola.



Since, all events on a hyperbola have the same invariant parameter $\Delta s = a$,

→ Consider in O :

$$\Delta x = 0, \Delta t = 1$$

$$\Rightarrow \Delta x^2 - \Delta t^2 = 1 = \underbrace{\Delta x'^2 - \Delta t'^2}_{\substack{\text{as measured} \\ \text{in } O'}}$$

$$\Rightarrow -\Delta t^2 = -\Delta t'^2 \left(1 - \left(\frac{\Delta x'}{\Delta t'}\right)^2\right)$$

↳ v as measured in O'

$$\Rightarrow \boxed{\Delta t' = \frac{\Delta t}{\sqrt{1-v^2}}} \leftarrow \text{Time dilation.}$$

H/W: Do a Her calculation when $\Delta t = 0$ & $\Delta x \neq 0$

↓
Length contraction

$$\Delta x' = \Delta x \sqrt{1-v^2}$$

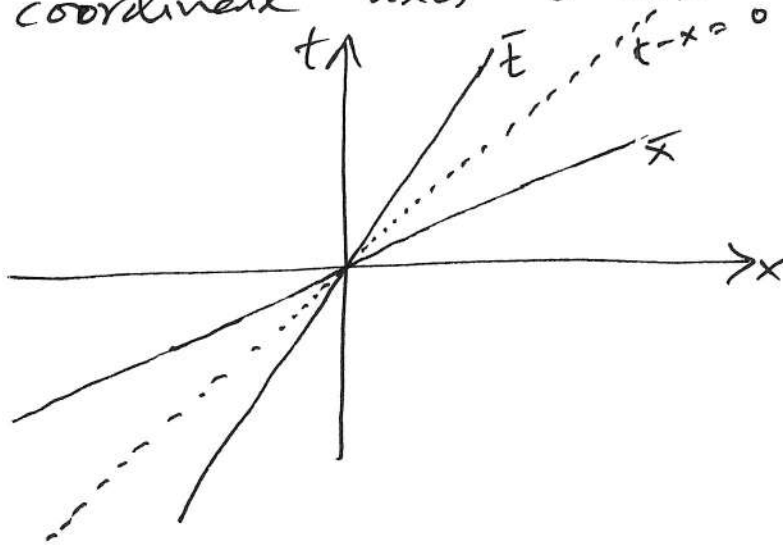
→ Proper time. $\Delta \tau^2 \equiv -\Delta s^2$

$$\Rightarrow \boxed{\Delta \tau \equiv \sqrt{\Delta t^2 - \Delta x^2 - \Delta y^2 - \Delta z^2}}$$

→ Paradoxes discussion from Sec. 1.8 Schutz, Schwartz.

The Lorentz Transformations:

→ From our geometrical understanding, the coordinate axes \bar{t} & \bar{x} wrt t & x :



From geometry:

Eq. of \bar{t} : ~~$t - vx = 0$~~

$vt - x = 0$

— (1) $v \leq c$

Eq. of \bar{x} :

$-t + vx = 0$

— (2)

→ Attempting to construct a linear transformation:

$\bar{t} = \alpha(t - vx)$; $\bar{x} = \sigma(x - vt)$

From ~~$\Delta x^2 + \Delta t^2 = 0$~~ $\Delta x^2 - \Delta t^2 = \Delta x'^2 - \Delta t'^2$

↳
$$\bar{t} = \frac{t - vx}{\sqrt{1 - v^2}} ; \bar{x} = \frac{x - vt}{\sqrt{1 - v^2}}$$

↑ Lorentz transformation.

H/W: What if transformation is linear?