

Towards GTR

So far :

→ Due to equivalence principle, the geometry of space in presence of gravity cannot be Euclidean !! ← Non-flat!!

→ Regardless of gravity, the geometry of space and time is not Galilean / Euclidean due to the modified principle of relativity [special relativity!]

↓

- Time is not universal
- speed of light is universal.

∴ a new geometry of "spacetime" where time & space are treated on equal footing.

↓

Minkowski spacetime

→ Now, getting back to the Einstein's realization on equivalence principle:

Q How to describe this "non-flat"/curved geometry of spacetime under gravity that is also consistent with special relativity?

Ans: GTR! **BUT** we first develop a formalism to consistently describe STR geometry.

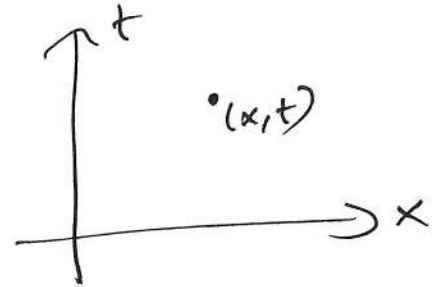
Review of Tensors

The metric tensor:

The Einstein notations:

→ We recall that in Minkowski geometry, time is also a coordinate label for an event.

→ There fore, an intuitive notation for an event:



$u^\alpha = (t, x, y, z)$ → Similar to a 3-vector
 $u^i = (x, y, z) \equiv \vec{u}$

→ Clearly, to find the norm, one has to compute $\sim \sum_\alpha u^\alpha u^\alpha$ (llor to $u^i u^i$ in Galilean Euclidean geom.)

→ Let's assume, that the repetition of indices implies a summation.

$$\Rightarrow \sum_\alpha u^\alpha u^\alpha \longrightarrow u^\alpha u^\alpha$$

→ An immediate observation:

How would one obtain an invariant distance (llor to 3D displacement: $u^i u^i$)?

Using the assignment for u^α above:

$$u^\alpha u^\alpha = u^\alpha \equiv t \hat{t} + x \hat{x} + y \hat{y} + z \hat{z}$$

→ our hypothesis*

$\Rightarrow u^\alpha u^\alpha = t^2 + x^2 + y^2 + z^2$ assuming that $\hat{t} \cdot \hat{t} = 1 = \hat{x} \cdot \hat{x} = \hat{y} \cdot \hat{y} = \hat{z} \cdot \hat{z}$

WRONG!!

\therefore We know, $-t^2 + x^2 + y^2 + z^2 = s^2$ in Minkowski geometry.

$\Rightarrow s^2 = \cancel{[\eta] u^\alpha u^\alpha} [\eta] [u] [u] = -t^2 + x^2 + y^2 + z^2$
 $\equiv \eta_{\alpha\beta} u^\alpha u^\beta$

with $\eta_{\alpha\beta} = \begin{pmatrix} -1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix} \leftarrow \text{Metric}$

→ Definition of a tensor: ** Skipping technicalities for a later course. (not within current scope)*

A tensor of type $\begin{pmatrix} 0 \\ N \end{pmatrix}$ is a function of N vectors into the real numbers, which is linear in each of its N arguments.

eg.: $g(\vec{A}, \vec{B}) \equiv \vec{A} \cdot \vec{B} \leftarrow \begin{pmatrix} 0 \\ 2 \end{pmatrix}$ tensor.
 $f(t, x, y, z) \leftarrow \begin{pmatrix} 0 \\ 0 \end{pmatrix}$ tensor.

* For some unit vectors $\vec{e}_{\alpha, \beta}$:

$g(\vec{e}_\alpha, \vec{e}_\beta) = \vec{e}_\alpha \cdot \vec{e}_\beta \equiv \eta_{\alpha\beta} \leftarrow \boxed{\text{Metric tensor}}$

→ If p_α is a tensor

$$p_\alpha \equiv \tilde{p}(\vec{e}_\alpha) \leftarrow \begin{pmatrix} 0 \\ 1 \end{pmatrix} \text{ tensor}$$

where \vec{e}_α is a unit vector.

1-form

* Note that the position of index is as subscript for 1-form.

$$\Rightarrow \begin{pmatrix} 0 \\ N \end{pmatrix} \text{ tensor: } p_{\alpha\beta\dots\gamma} \equiv \tilde{p}(\vec{e}_\alpha, \vec{e}_\beta, \dots, \vec{e}_\gamma)$$

N-form

Q Based on this, what can you say about the decomposition a rank-2 antisymmetric field $B_{\mu\nu}$ into scalar/vector/tensor components?

→ Now, for a vector $\vec{A} \equiv A^\alpha \vec{e}_\alpha$

$$\Rightarrow \tilde{p}(\vec{A}) = \tilde{p}(A^\alpha \vec{e}_\alpha) = A^\alpha \tilde{p}(\vec{e}_\alpha)$$

$$\Rightarrow \boxed{\tilde{p}(\vec{A}) = A^\alpha p_\alpha} \leftarrow \text{1-form for a gen. vector } \vec{A}.$$

⇒ A 1-form is different from a ⁴⁻vector!

Q Relation between a 1-form (A_α ?) and corresponding vector A^α ?

→ Since the metric is a mapping given by:

$$g(\vec{A}, \vec{B}) \equiv \vec{A} \cdot \vec{B}$$

~~For a 1-form:~~

$$\tilde{p}(\vec{A})$$

Suppose $g(\vec{V}, \vec{A}) = \vec{V} \cdot \vec{A}$

Let's define a 1-form \tilde{V} :

$$\boxed{\tilde{V}(\vec{A}) \equiv \vec{V} \cdot \vec{A}}$$

($\because \vec{V} \cdot$ is permanent but \vec{A} is the argument
 \downarrow
 satisfies the $\binom{0}{1}$ def.)

⇒ The components of such a 1-form:

$$\begin{aligned} V_\alpha &\equiv \tilde{V}(\vec{e}_\alpha) = \vec{V} \cdot \vec{e}_\alpha = \vec{e}_\alpha \cdot \vec{V} \\ &= \vec{e}_\alpha \cdot V^\beta \vec{e}_\beta \end{aligned}$$

$$\Rightarrow V_\alpha = (\vec{e}_\alpha \cdot \vec{e}_\beta) V^\beta \equiv \delta_{\alpha\beta} V^\beta$$

$$\Rightarrow \boxed{V_\alpha \equiv \eta_{\alpha\beta} V^\beta} \leftarrow \text{Finally looks familiar!!}$$

In physics texts:

- $V_\alpha \rightarrow$ covariant 4-vector \leftarrow "1-form"
- $V^\alpha \rightarrow$ contravariant 4-vector.

*Note: Important to distinguish 1-form and vector due to the difference in gradient components or just components.

For example:

Consider a vector $V^\alpha = (a, b, c, d)$

the difference b/w form & vec. \Rightarrow ~~$V_\beta = (-a, b, c, d)$~~

$$V_\beta = \eta_{\beta\alpha} V^\alpha = \begin{pmatrix} -1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix} \begin{pmatrix} a \\ b \\ c \\ d \end{pmatrix}$$

$$= \boxed{(-a, b, c, d)}$$

1ly

$$\begin{aligned} \tilde{d}\phi &\rightarrow \left(\frac{\partial\phi}{\partial t}, \frac{\partial\phi}{\partial x}, \dots \right) \\ d\phi &\rightarrow \left(-\frac{\partial\phi}{\partial t}, \frac{\partial\phi}{\partial x}, \dots \right) \end{aligned} \quad \left. \vphantom{\begin{aligned} \tilde{d}\phi \\ d\phi \end{aligned}} \right\} \text{how?}$$

\rightarrow To understand the above, we will now review the tensor calculus.

* If an inverse of the metric exists, i.e.:

$$(\eta_{\alpha\beta})^{-1} = \eta^{\alpha\beta} \quad \leftarrow \text{Non-vanishing determinant.}$$

then, $\boxed{A^\alpha = \eta^{\alpha\beta} A_\beta}$

→ Magnitude of a vector & 1-form:

$$\vec{p}^2 = \eta_{\alpha\beta} p^\alpha p^\beta = \vec{p}^2$$

~~$$\vec{p}^2 = \eta_{\alpha\beta}$$~~

$$\vec{p}^2 = \eta_{\alpha\beta} p^\alpha p^\beta$$

$$= \eta_{\alpha\beta} (\eta^{\alpha\mu} p_\mu) (\eta^{\beta\nu} p_\nu)$$

$$= \eta^{\alpha\mu} p_\mu (\eta_{\alpha\beta} \eta^{\beta\nu}) p_\nu$$

↓ Inverse property.

$$= \eta^{\alpha\mu} p_\mu \delta_\alpha^\nu p_\nu$$

$$\Rightarrow \boxed{\vec{p}^2 = \eta^{\alpha\mu} p_\alpha p_\mu = \vec{p}^2}$$

For general vec. : $p_\mu q^\mu \rightarrow$ scalar product.

→ Raising & lowering indices:

Easy to infer that going from co- to contra- is through the use of metric.

$$p_\mu = \eta_{\mu\nu} p^\nu \quad \& \quad p^\mu = \eta^{\mu\nu} p_\nu$$

Q For the flat Minkowski metric $\eta_{\alpha\beta}$, what is η^α_β ?

Differentiation & Calculus of tensors -
(without the baggage of mathematics!)

$$\frac{\partial p^\nu}{\partial x^\alpha} \equiv p^\nu_{,\alpha}$$

$$\frac{\partial p_\nu}{\partial x^\alpha} \equiv p_{\nu,\alpha} = \frac{\partial}{\partial x^\alpha} (\eta_{\nu\mu} p^\mu) = \eta_{\nu\mu} \frac{\partial p^\mu}{\partial x^\alpha}$$

↓
iff $\eta_{\mu\nu} \rightarrow$ Minkowski

In general curved manifold:

$$\frac{\partial p_\nu}{\partial x^\alpha} = \frac{\partial}{\partial x^\alpha} (g_{\nu\mu} p^\mu) = g_{\mu\nu,\alpha} p^\mu + g_{\mu\nu} p^\mu_{,\alpha}$$

Now, Recall - Back to Physics:

We have figured out the mathematical framework to describe Minkowski spacetime of STR, using tensors (3+1D) & Minkowski metric.

→ Now, we must introduce curvature in STR that is consistent with equivalence principle.

HW [Read Chap 5.1, Schutz]

→ Local Lorentz frame and equivalence principle.